

Od Kopernika až po Keplerove zákony

Keplerove zákony

Newtonov gravitačný zákon

Veličiny charakterizujúce gravitačné pole

intenzita, potenciálna energia, potenciál

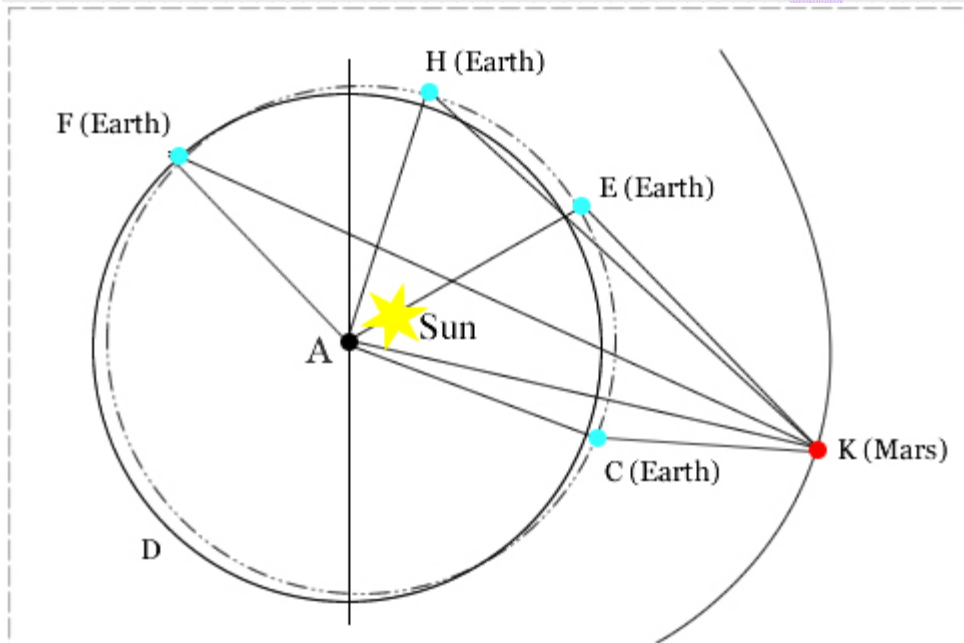
Pohyb telies v gravitačnom poli Slnka

1. a 2. kozmická rýchlosť

geocentrický názor

Ptolemaios + Aristoteles

Tycho Brahe (1546-1601)
Planéty sa pohybujú okolo Slnka, ale spolu so Slnkom rotujú okolo Zeme, ktorá je stredom vesmíru



heliocentrický názor

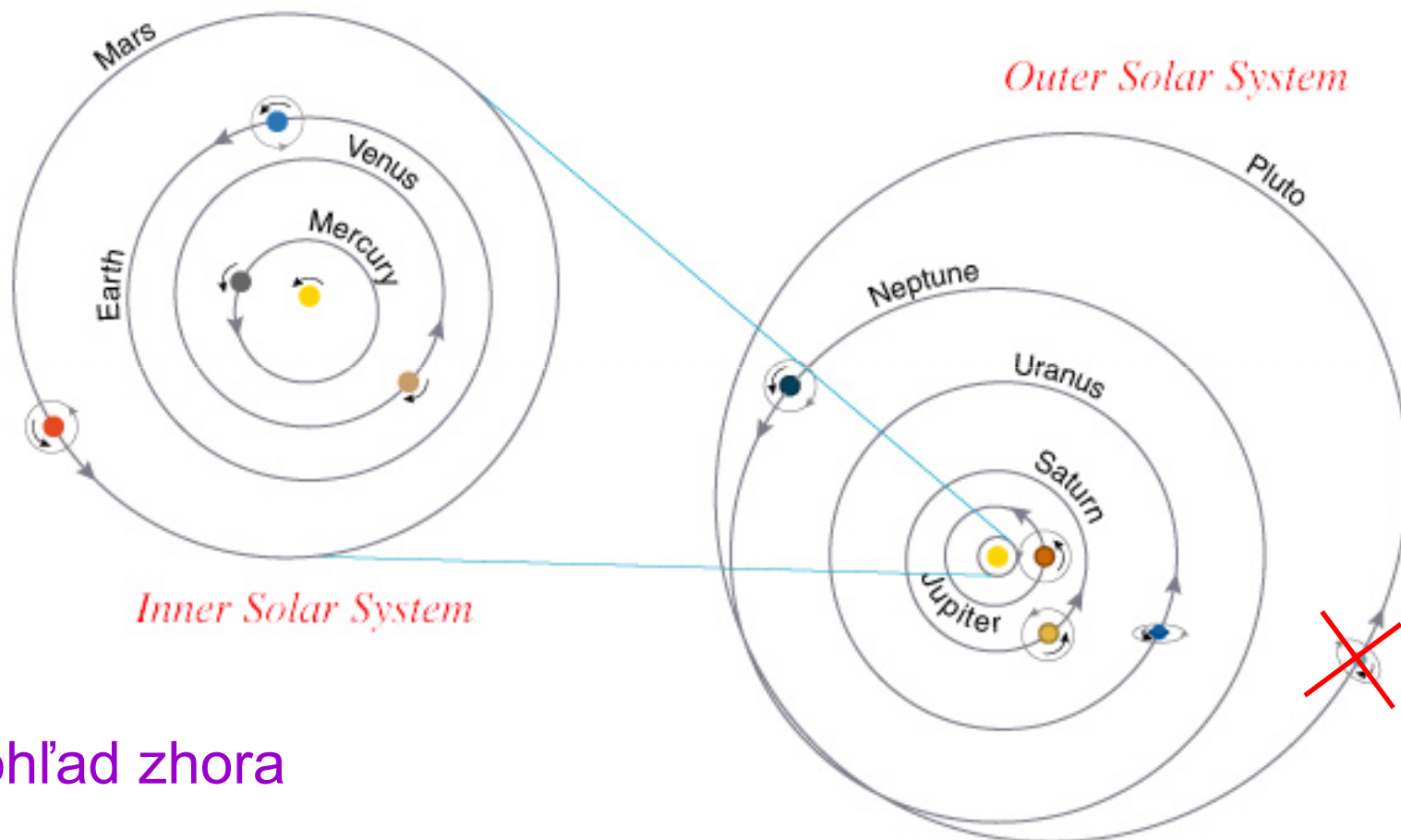
Nicolaus Copernicus (1473-1543)
Zem sa točí okolo osi aj okolo Slnka, planéty sa pohybujú okolo Slnka rovnakým smerom.

Giordano Bruno (1548-1600)
Kopernikus+ navyše ani Slnko nie je stred vesmíru ale len jedna z hviezd

Johannes Kepler (1571-1630)
zhrnul pozorovania TB a NC a **sformuloval Keplerove zákony**, ako prvý dobre vysvetlil nezrovnalosti v pozorovaniach tým, že zaviedol eliptické dráhy

Galileo Galilei (1564-1642)
zostrojil ďalekohľady s 30x zväčš., podporil ešte viac heliocentrizmus

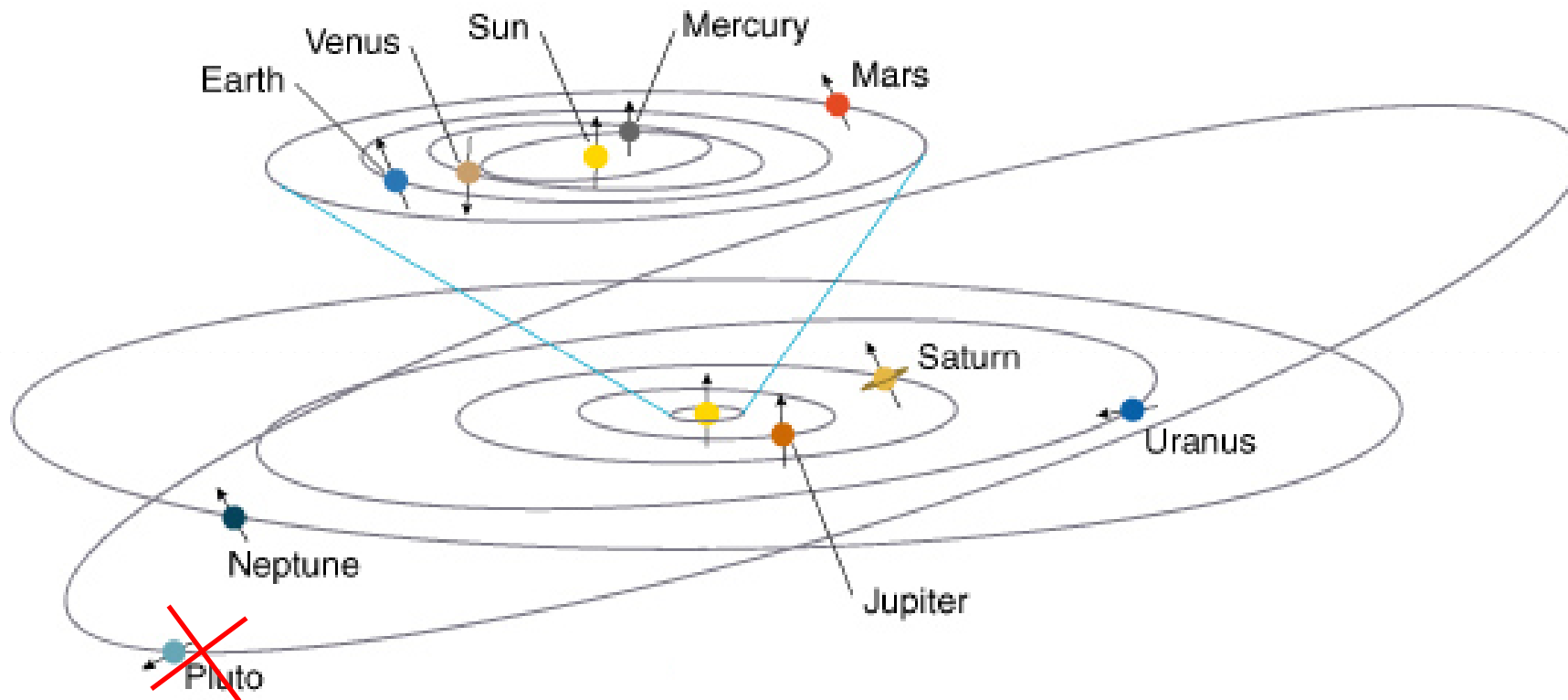
Všetky planéty obiehajú okolo Slnka po eliptických dráhach, pričom Slnko je v ich spoločnom ohnisku



pohľad zhora

Pozn. elipsa má len málo odlišné dĺžky polosí (napr. Zem je v Apheliu 4. júla – 152.000.000km a v Periheliu 3. januára – 149.600.000km)

Pluto bolo v roku 2006 vyradené zo zoznamu planét Slnčnej sústavy!

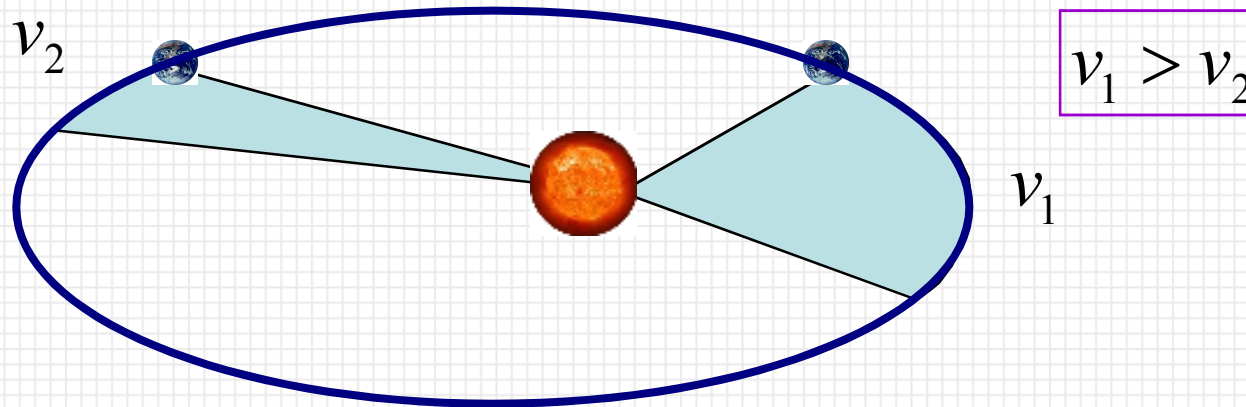


pohľad v rovine obežných dráh

	Perihelion [AU]	Aphelion [AU]	Period [day]	Mass [10^{24} kg]
Mercury	0.31	0.47	88,96	0,335
Venus	0.72	0.73	224,54	4,89
Earth	0.98	1.02	365,25	5,98
Mars	1.38	1.67	687,50	0,646
Jupiter	4.95	5.45	4328,7	1900
Saturn	9.02	10.0	10822	569
Uranus	18.3	20.1	30555	87,3
Neptune	30.0	30.3	60417	103
Pluto	29.7	49.9	90509	5,4

1AU... priemerná vzdialenosť Zem-Slnko 152.000.000km

2. Keplerov zákon



Sprievodič spájajúci Slnko s planétou opisuje za rovnaký čas rovnaké plochy

3. Keplerov zákon

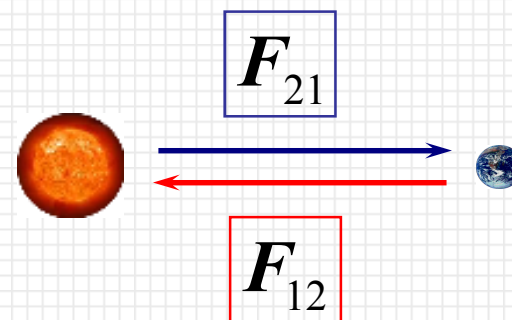
$$\frac{T_1^2}{r_1^3} = \frac{T_2^2}{r_2^3}$$

$$\frac{T^2}{r^3} = \textit{konšt.}$$

(pomocou 3. Keplerovho zákona)

$$a_n = \frac{v^2}{r} = \omega^2 r = \left(\frac{2\pi}{T} \right)^2 r = r \frac{4\pi^2}{T^2} = 5,929 \cdot 10^{-3} \text{ ms}^{-2}$$

$$T^2 = 4\pi^2 \frac{r}{a} \xrightarrow{\text{3.K.z.}} 4\pi^2 \frac{r}{ar^3} = \text{konšt.}$$



$$a = k \frac{1}{r^2}$$

$$F_{12} = ma = k \frac{m}{r^2}$$

$$k = M\kappa$$

$$F_{21} = Ma' = k' \frac{M}{r^2}$$

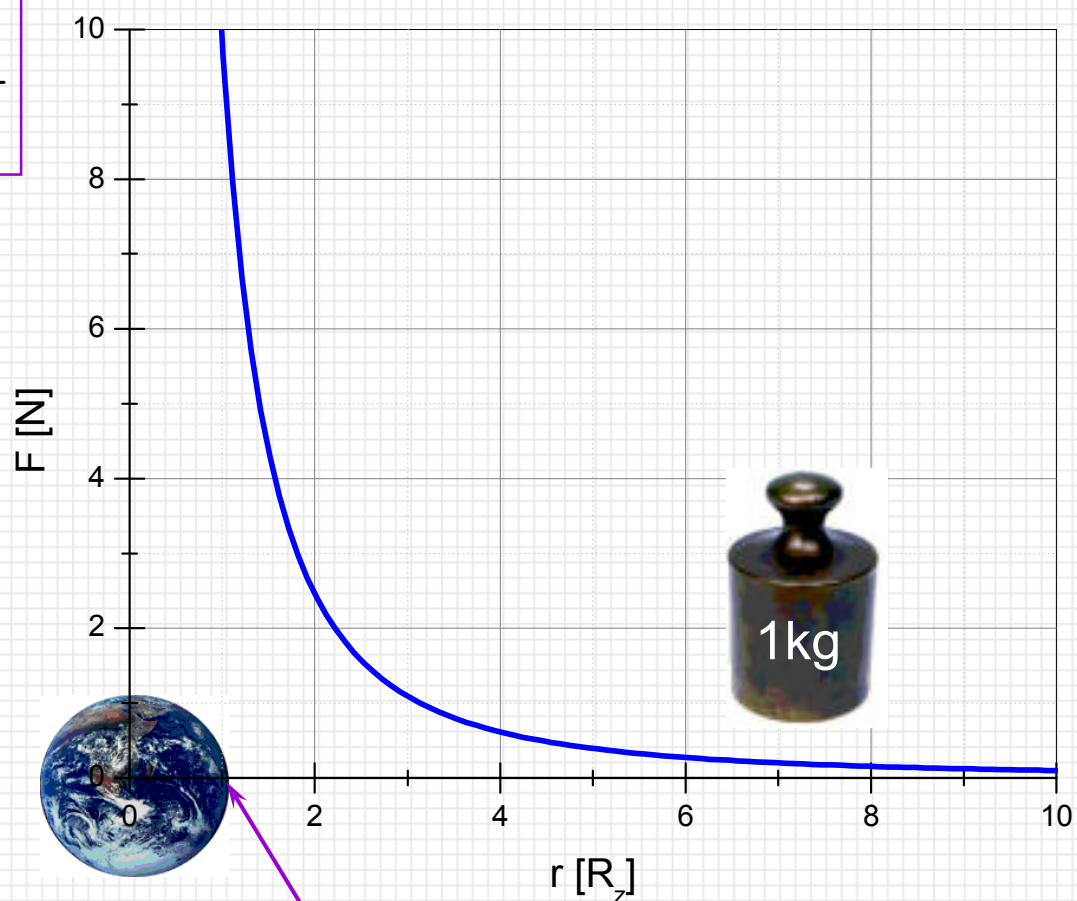
$$k' = m\kappa$$

$$F_{12} = -\kappa \frac{Mm}{r^3} \mathbf{r}_{12}$$

$$\kappa = 6,671 \cdot 10^{-11} \text{ Nm}^2 \text{ kg}^{-2}$$

Závislosť veľkosti gravitačnej sily od vzdialenosti od Zeme na hm. b. 1kg

$$F = \kappa \frac{Mm}{r^2}$$



povrch Zeme

Porovnanie veľkosti gravitačnej sily na povrchu Mesiaca a Zeme

$$m = 1\text{kg}$$



$$F = \kappa \frac{Mm}{R^2}$$



$$M_M = 7,349 \cdot 10^{22} \text{ kg}$$

$$R_M = 1,738 \cdot 10^6 \text{ m}$$

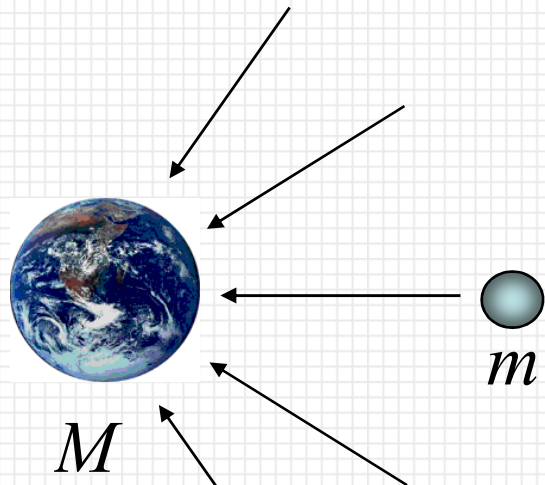
$$F_M = 1,62 \text{ N}$$

$$M_Z = 5,974 \cdot 10^{24} \text{ kg}$$

$$R_Z = 6,371 \cdot 10^6 \text{ m}$$

$$F_Z = 9,81 \text{ N}$$

$$\frac{F_Z}{F_M} \cong 6$$



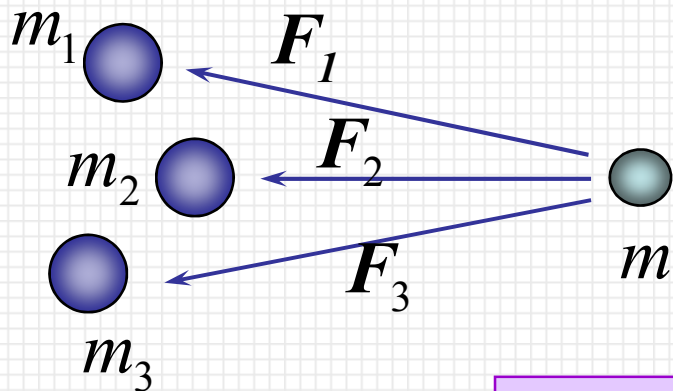
$$\mathbf{K} = \frac{\mathbf{F}}{m}$$



$$\mathbf{K} = -\kappa \frac{M}{r^3} \mathbf{r}$$

$$\mathbf{F} = -\kappa \frac{Mm}{r^3} \mathbf{r}$$

Intenzita v okolí sústavy hm. b. a telesa



$$\mathbf{K} = \frac{\mathbf{F}}{m} = \frac{\mathbf{F}_1 + \mathbf{F}_2 + \dots + \mathbf{F}_n}{m}$$

$$= \frac{\mathbf{F}_1}{m} + \frac{\mathbf{F}_2}{m} + \dots + \frac{\mathbf{F}_n}{m} = \mathbf{K}_1 + \mathbf{K}_2 + \dots + \mathbf{K}_n$$

$$\mathbf{K} = \sum_{i=1}^n \mathbf{K}_i$$

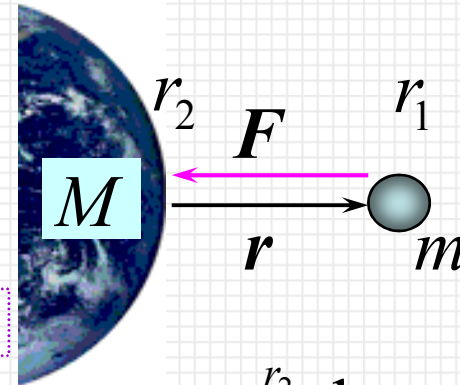
$$= -\kappa \sum \frac{m_i}{r_i^3} \mathbf{r}_i$$

$$= -\kappa \int \frac{dm}{r^3} \mathbf{r}$$

$$= -\kappa \int \frac{\mathbf{r}}{r^3} \rho d\tau$$

$$E_p = ?$$

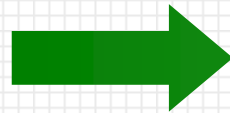
$$E_p = -A$$



z Newtonovho gravitačného zákona

$$E_p = - \int_{r_1}^{r_2} \mathbf{F} \cdot d\mathbf{r} = \kappa m M \int_{r_1}^{r_2} \frac{\cancel{\mathbf{r}} \cdot d\mathbf{r}}{r^{\cancel{3}2}} = \kappa m M \int_{r_1}^{r_2} \frac{dr}{r^2}$$

$$E_p = \kappa m M \left[-\frac{1}{r} \right]_{r_1}^{r_2}$$



$$E_p = -\kappa m M \left(\frac{1}{r_2} - \frac{1}{r_1} \right)$$

$$A = E_p = -\kappa m M \left(\frac{1}{r_2} - \frac{1}{r_1} \right) = -\frac{\kappa m M}{r} \Big|_{r_1 \rightarrow \infty} + 0$$

$$E_p = -\kappa m M \frac{1}{r}$$

$$V = \frac{E_p}{m} \quad \longrightarrow \quad V = -\kappa \frac{M}{r}$$

$$[V] = \frac{\text{J}}{\text{kg}} = \frac{\text{kgm}^2\text{s}^{-2}}{\text{kg}} = \text{m}^2\text{s}^{-2}$$

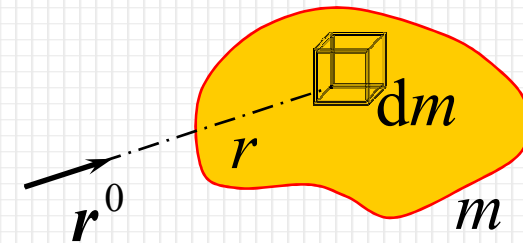


pre sústavu hm. b. plátí:

$$V = -\kappa \sum_i \frac{m_i}{r_i}$$

pre teleso so spojitou
rozloženou hmotnosťou platí:

$$V = -\kappa \int \frac{dm}{r}$$



$$F = mK \longleftrightarrow F_g = mg$$

~~$$mK = mg$$~~

$$K = g$$

$$K = -\kappa \frac{M}{r^3} \mathbf{r} = \mathbf{g}$$

$$g_h = \kappa \frac{M}{(R+h)^2}$$

vo výške h nad zemským povrchom

$$g_0 = \kappa \frac{M}{R^2}$$

na povrchu Zeme

$$g_{Zem} = -6.671 \cdot 10^{-11} \text{ Nkg}^{-2} \text{ m}^2 \frac{5.974 \cdot 10^{24} \text{ kg}}{(6.371 \cdot 10^6 \text{ m})^2} = 9.818 \text{ ms}^{-2}$$

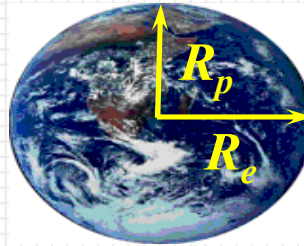
$$g_{Mesiac} = 1.623 \text{ ms}^{-2}$$

$$\frac{g_{Zem}}{g_{Mesiac}} = 6.05$$

$$M = 5,974 \cdot 10^{24} \text{ kg}$$

$$R_e = 6,38 \cdot 10^6 \text{ m}$$

$$R_p = 6,36 \cdot 10^6 \text{ m}$$



$$g_p = \kappa \frac{M}{R_p^2}$$

$$g_p = 6,671 \cdot 10^{-11} \text{ Nm}^2 \text{ kg}^{-2} \frac{5,974 \cdot 10^{24} \text{ kg}}{(6,36 \cdot 10^6 \text{ m})^2}$$

$$g_p = 9,86 \text{ ms}^{-2}$$

$$F_{op} = 0$$

? odstredivá sila

$$g'_p = 9,86 \text{ ms}^{-2}$$

$$g_e = \kappa \frac{M}{R_e^2}$$

$$g_e = 6,671 \cdot 10^{-11} \text{ Nm}^2 \text{ kg}^{-2} \frac{5,974 \cdot 10^{24} \text{ kg}}{(6,38 \cdot 10^6 \text{ m})^2}$$

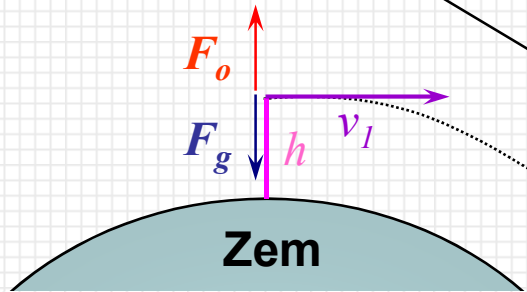
$$g_e = 9,80 \text{ ms}^{-2}$$

$$F_{oe} = \frac{4\pi^2 R_e}{T^2}$$

$$g'_e = g_e - \frac{4\pi^2 R_e}{T^2} = 9,80 \text{ ms}^{-2} - \frac{39,48 \cdot 6,38 \cdot 10^6 \text{ m}}{86400^2 \text{ s}^2} = 9,77 \text{ ms}^{-2}$$

Kozmické rýchlosti

- 1. kozm. rýchlosť – obežnica Zeme
- 2. kozm. rýchlosť – opustenie gr. poľa Zeme
- 3. kozm. rýchlosť – opustenie Slnčej sústavy



$$F_g = F_o$$

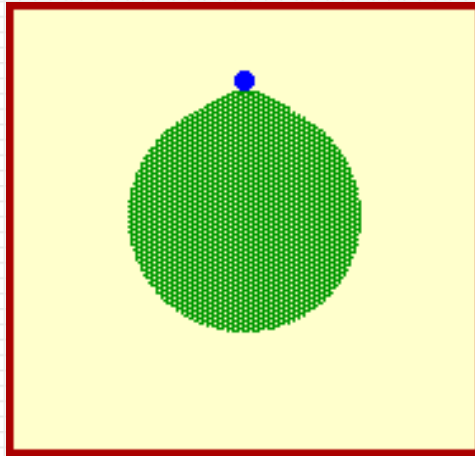
$$g = \kappa \frac{M}{(R+h)^2} \longleftrightarrow g_0 = -\kappa \frac{M}{R^2}$$

$$g = g_0 \frac{R^2}{(R+h)^2}$$

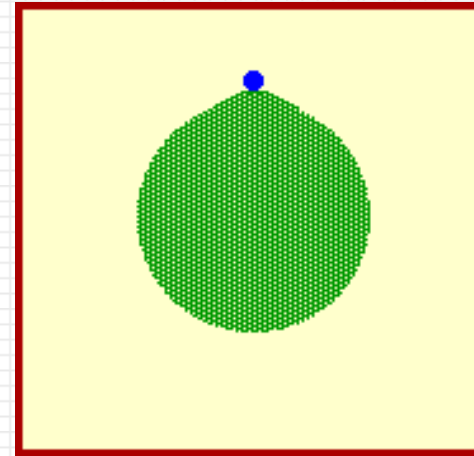
$$mg = m \frac{v_1^2}{R+h}$$

$$g_0 \frac{R^2}{(R+h)^2} = \frac{v_1^2}{R+h}$$

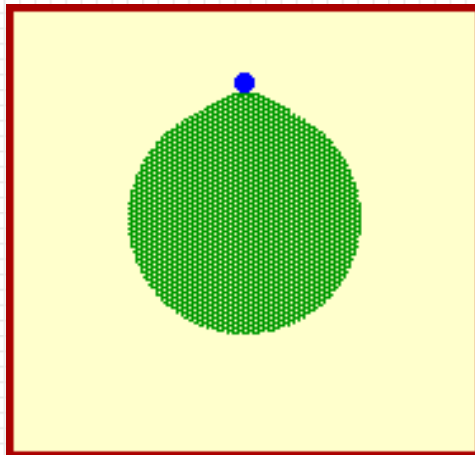
$$v_1 = R \sqrt{\frac{g_0}{R+h}} \quad \text{na povrchu Zeme } h=0 \quad \longrightarrow \quad v_1 = \sqrt{Rg_0} = 7.912 \text{ km s}^{-1}$$



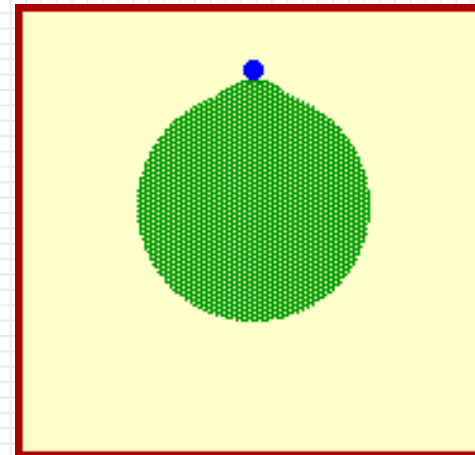
$$v = 4 \text{ km s}^{-1}$$



$$v = 6 \text{ km s}^{-1}$$



$$v = 8 \text{ km s}^{-1}$$

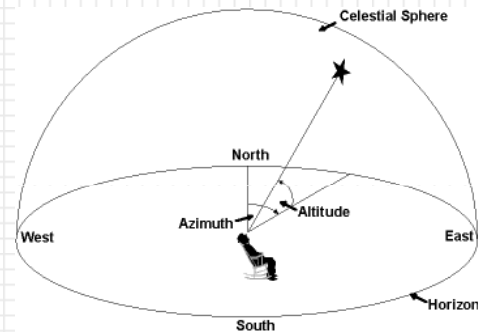


$$v > 8 \text{ km s}^{-1}$$

Date	Mag	Starts			Max. altitude			Ends		
		Time	Alt.	Az.	Time	Alt.	Az.	Time	Alt.	Az.
17 Mar	-1.9	19:27:42	10	W	19:30:17	30	SSW	19:31:05	25	S
18 Mar	-3.1	18:16:57	10	WNW	18:19:49	66	SSW	18:22:40	10	ESE
18 Mar	-0.6	19:53:10	10	WSW	19:54:34	13	SW	19:54:59	13	SSW
19 Mar	-1.7	18:41:39	10	W	18:44:13	29	SSW	18:46:46	10	SSE
20 Mar	-0.3	19:07:05	10	WSW	19:08:24	12	SW	19:09:42	10	SSW
22 Mar	-0.1	18:20:51	10	WSW	18:22:02	12	SW	18:23:13	10	SSW

location: Zilina, 49.2170°N, 18.7330°E

N (0°), NNE (22.5°), NE (45°), ENE (67.5°), E (90°), ESE (112.5°), SE (135°), SSE (157.5°), S (180°), SSW (202.5°), SW (225°), WSW (247.5°), W (270°), WNW (292.5°), NW (315°), NNW (337.5°)



<http://heavens-above.com>

S119E:09T62

Magnitude

Sun (about 400 000 times brighter than full Moon!)	-26.7
Full Moon	-12.7
Brightest Iridium flares	-8
Venus (at brightest)	-4.4
International Space Station	-2
Sirius (brightest star)	-1.44
Limit of human eye	+6 to +7
Limit of 10x50 binoculars	+9
Pluto	+14
Limit of Hubble Space Telescope	+30

